

# Minor Constituents of Mars Atmosphere

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- Martian atmosphere consists of CO<sub>2</sub> and its products, H<sub>2</sub>O and its products, N<sub>2</sub> and noble gases
- Complicated seasonal variations induced by obliquity of 24° and the elliptic orbit result in a great variety of conditions from T = 145 K at the winter polar regions to 300 K at low latitudes near perihelion

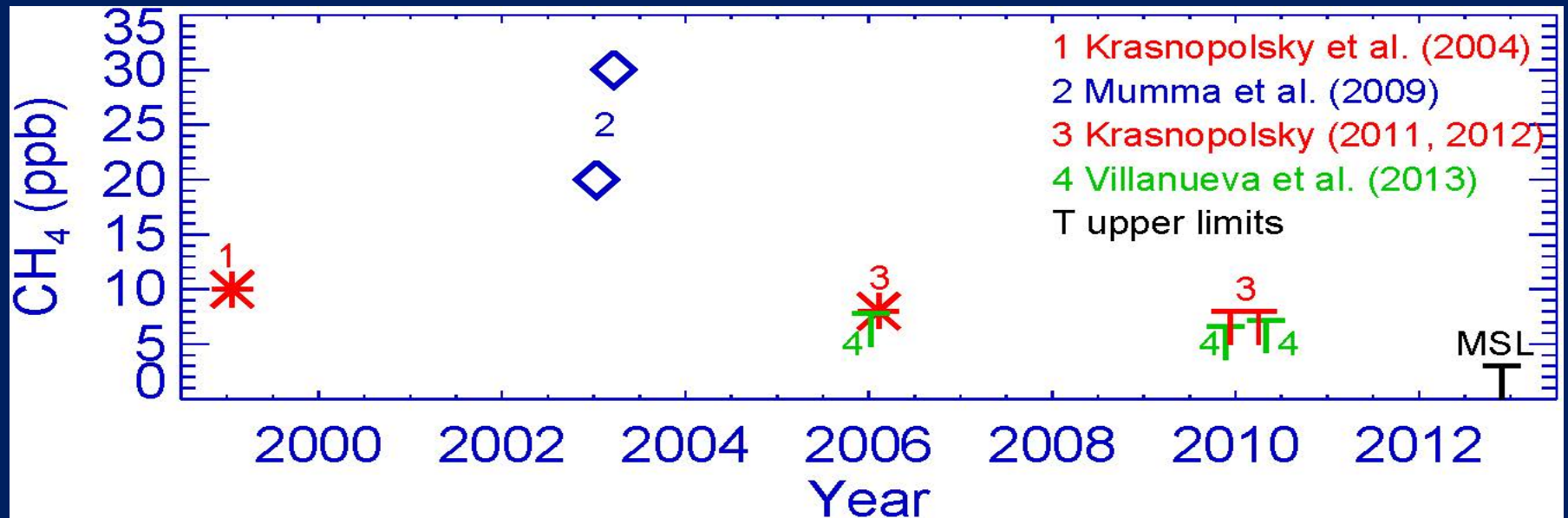
- CO<sub>2</sub> and products: CO, O<sub>2</sub>, O<sub>2</sub>(1Δg), O<sub>3</sub>, O
- CO: observations ~1000 ppm, models ~120 ppm  
 CO production: CO<sub>2</sub> + hv (7x10<sup>11</sup> cm<sup>-2</sup> s<sup>-1</sup>)  
 CO loss: CO + OH → CO<sub>2</sub> + H,  $k = 1.5 \times 10^{-13}$
- Variations of CO and O<sub>2</sub> are induced by condensation and sublimation of CO<sub>2</sub> in the polar and subpolar regions
- O<sub>2</sub>(1Δg): product of O<sub>3</sub> photolysis, dayglow at 1.27 μm, tracer of photochemistry, polar nightglow, nightglow at low latitudes(?)

H<sub>2</sub>O and products: H, H<sub>2</sub>, H<sub>2</sub>O<sub>2</sub>, OH,  
HO<sub>2</sub>

- H<sub>2</sub>O vertical profiles (continuation of SPICAM-IR)
- H<sub>2</sub>O<sub>2</sub> abundances are diagnostics of variations of odd hydrogen chemistry on Mars
- Ground-based TEXES observations (8 μm):  
0-40 ppb (Encrenaz et al. 2012), within model predictions (Lefevre et al. 2008, Krasnopolsky 2009)

- **N<sub>2</sub>, Ar, He, Ne, Kr, Xe** : unmeasurable
- NO  $\approx$  0.3 ppb (models); less than 1.7 ppb (5.3  $\mu$ m, TEXES, Kr2006); TIRVIM?
- NO<sub>2</sub>  $\approx$  0.1 ppb (models); less than 10 ppb (M9, 6  $\mu$ m, Maguire 1977); TIRVIM? 3.45  $\mu$ m MIR?
- N<sub>2</sub>O  $\approx$  1 ppt (model); <100 ppb (M9, 8  $\mu$ m, M77), <70 ppb (NIRSPEC, 3  $\mu$ m, Villanueva13)

# Ground-based and MSL observations of methane on Mars



MSL observations of methane neither contradict nor support the ground-based observations

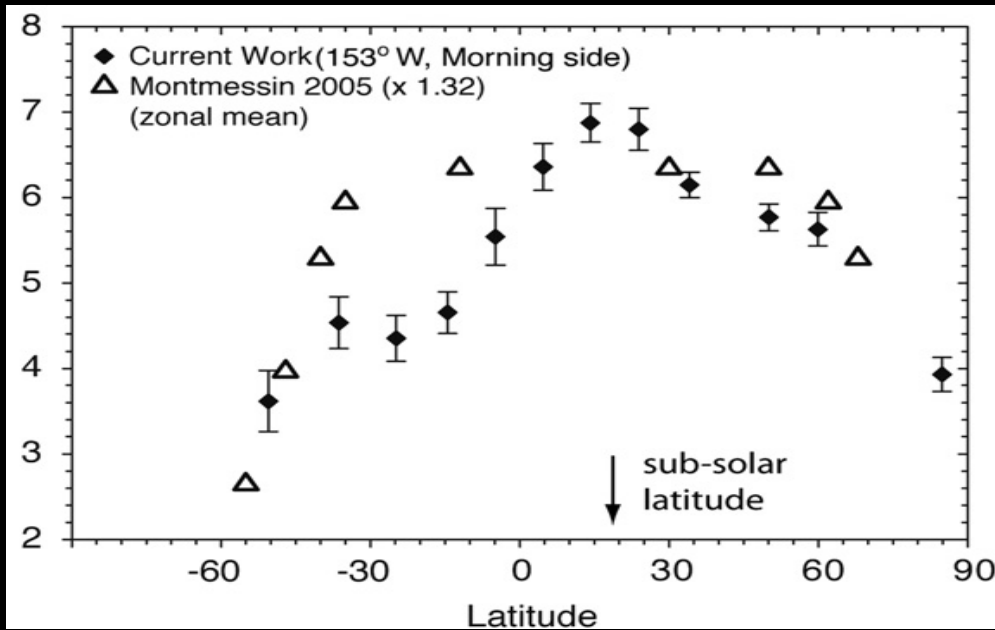
C<sub>2</sub>H<sub>6</sub> < 0.2 ppb (3.3  $\mu$ m, Kr2012, V2013)

C<sub>2</sub>H<sub>6</sub> >> other hydrocarbons (expected)

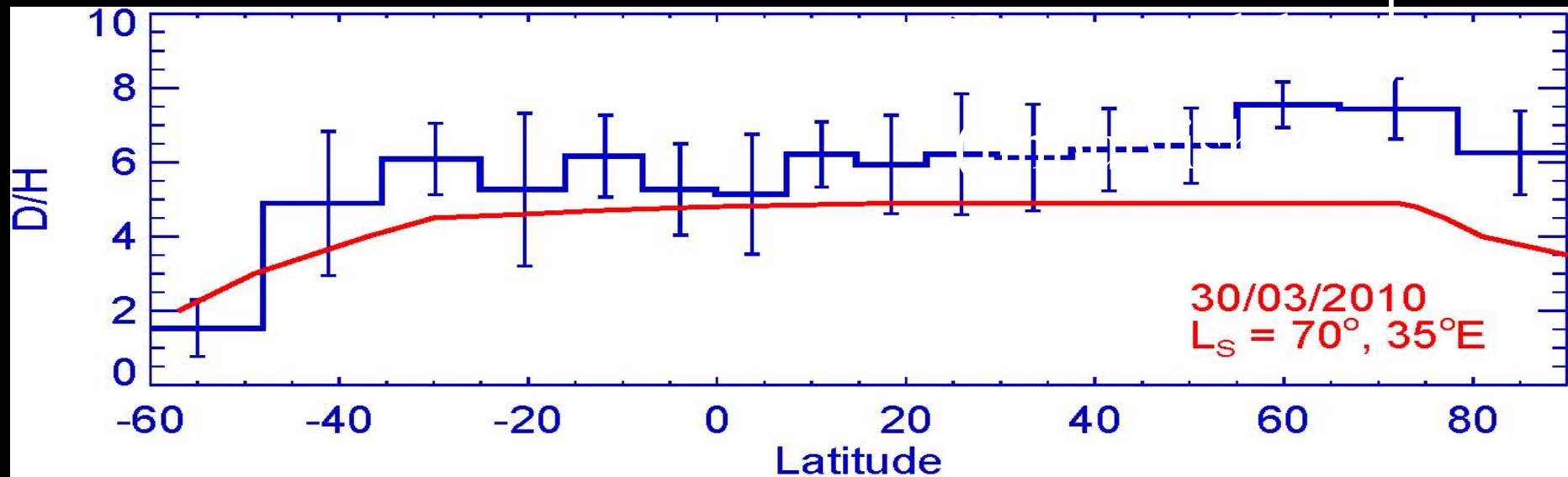
# Other important species

- $\text{SO}_2 < 0.3$  ppb (TEXES,  $8 \mu\text{m}$ , Encrenaz et al. 2011, Krasnopolsky 2012)  
Martian outgassing is smaller than that on Earth by a factor of more than 2000
- $\text{HCl} < 0.2$  ppb (Herschel, Hartogh et al. 2010);  
problem of perchlorates
- $\text{H}_2\text{CO} < 2$  ppb ( $3.6 \mu\text{m}$ , Krasnopolsky et al. 1997)  
The most abundant product of methane oxidation

# HDO/H<sub>2</sub>O observations and model



← LS = 50o, Novak et al. (2011)



# $^{13}\text{C}/^{12}\text{C}$ in methane

- If  $\delta^{13}\text{C} < -6\%$  then methane is of biologic origin, otherwise it is geological.
- Required uncertainty is  $\sim 1\%$ ,  $^{13}\text{CH}_4$  is less abundant than  $^{12}\text{CH}_4$  by a factor of  $\sim 100$ , and measurements of  $^{13}\text{C}/^{12}\text{C}$  require an increase in sensitivity by three orders of magnitude or more